Periodic orbit sum rules for billiards: accelerating cycle expansions

This article has been downloaded from IOPscience. Please scroll down to see the full text article.
1999 J. Phys. A: Math. Gen. 326757
(http://iopscience.iop.org/0305-4470/32/39/304)
View the table of contents for this issue, or go to the journal homepage for more

Download details:
IP Address: 171.66.16.111
The article was downloaded on 02/06/2010 at 07:44

Please note that terms and conditions apply.

# Periodic orbit sum rules for billiards: accelerating cycle expansions 

Sune F Nielsen $\dagger$, Per Dahlqvist $\dagger$ and Predrag Cvitanović $\ddagger$<br>$\dagger$ Royal Institute of Technology, S-100 44 Stockholm, Sweden<br>$\ddagger$ Department of Physics and Astronomy, Northwestern University 2145 Sheridan Road,<br>Evanston, IL 60208-3112, USA

Received 21 December 1998, in final form 2 July 1999


#### Abstract

We show that the periodic orbit sums for two-dimensional billiards satisfy an infinity of exact sum rules. We demonstrate their utility by using the flow conservation sum rule to accelerate the convergence of cycle expansions for the overlapping three-disc billiard. The effectiveness of the approach is studied by applying the method on averages, known explicitly by other sum rules. The method is then applied to the Lyapunov exponent.


## 1. Introduction

Periodic orbit theory is a powerful tool for the description of chaotic dynamical systems [1-3]. However, as one is dealing with infinities of cycles, the formal theory is not meaningful unless supplemented by a theory of convergence of cycle expansions. For nice hyperbolic systems, the theory is well developed, and shows that exponentially many cycles suffice to estimate chaotic averages with super-exponential accuracy [5,6]. However, for generic dynamical systems with infinitely specified grammars and/or non-hyperbolic phase space regions, the convergence of the dynamical zeta functions and spectral determinants cycle expansions is less remarkable. The infinite symbolic dynamics problem is generic, and a variety of strategies for dealing with it have been proposed: stability truncations [7, 8], approximate partitions [9], noise regularization [10] and even abandoning the periodic theory altogether [11, 12].

Computation of periodic orbits for a given system often requires a considerable investment, as exhaustively locating the periodic orbits of increasing length for flows in higher dimensions can be a demanding chore. It is therefore essential that the information obtained be used as effectively as possible. Here, we propose a new, hybrid approach of combining cycle expansions with exact results for 'nearby' averages, based on the observation that the periodic orbit sums sometimes satisfy exact sum rules.

Studies of convergence of cycle expansions, such as comparisons [13] of truncation errors of the dimension and the topological entropy for the Hénon attractor, indicate strong correlations in truncation errors for different averages. We propose to turn these correlations in our favour, by using the error known exactly by a sum rule to improve the estimate for a nearby average for which no exact result exists. Billiards provide a convenient, physically motivated testing ground for this idea. The approach is inspired by formula (17) for mean free-flight time in billiards, so well known to the Russian school that it went unpublished for decades [14]. In this paper we show that billiards obey an infinity of exact periodic orbit sum rules, and indicate how such rules might be used to accelerate the convergence of cycle expansions.

The paper is organized as follows. Section 2 gives a brief summary of the theory of periodic orbit averaging. In section 3 we review the known exact sum rules for billiards, and then generalize them to an infinity of sum rules. In section 4 we present the conventional cycle expansion numerical results for our test system, the overlapping three-disc billiard. This system is hyperbolic and does not suffer from the intermittency effects that plague billiards such as the stadium and the Sinai billiards, but is still 'generic' in the sense that its symbolic dynamics is arbitrarily complicated. The numerics presented indicate the existence of an analytic region extending beyond the leading zero in the relevant zeta functions. Under this assumption, in section 5 and the appendix, we develop a method which utilizes the flow conservation sum rule to accelerate the convergence of cycle expansions, and apply the method to our test system.

## 2. Periodic orbit averaging

We start with a summary of the basic formulae of the periodic orbit theory-for details the reader can consult $[1,3]$.

A flow $\boldsymbol{x} \rightarrow \boldsymbol{f}^{t}(\boldsymbol{x}), x \in \mathcal{M}$, is a continuous mapping $\boldsymbol{f}^{t}: \mathcal{M} \rightarrow \mathcal{M}$ of the phase space $\mathcal{M}$ onto itself, parametrized by time $t$. On a suitably defined Poincaré surface of section $\mathcal{P}$, the dynamics is reduced to a return map

$$
\begin{equation*}
x \rightarrow f^{n}(x) \quad x \in \mathcal{P} \tag{1}
\end{equation*}
$$

where $n$ is the 'topological time', the number of times the trajectory returns to the surface of section.

A dynamical zeta function [4] associated with the flow $\boldsymbol{f}^{t}(\boldsymbol{x})$ is defined as the product over all prime cycles $p$

$$
\begin{equation*}
1 / \zeta(z, s, \beta)=\prod_{p}\left(1-t_{p}\right) \quad t_{p}=t_{p}(z, s, \beta)=\frac{1}{\left|\Lambda_{p}\right|} \mathrm{e}^{\beta A_{p}-s T_{p}} z^{n_{p}} \tag{2}
\end{equation*}
$$

where $T_{p}, n_{p}$ and $\Lambda_{p}$ are the period, topological length and stability of prime cycle $p$. Furthermore, $s$ is a variable dual to the time $t, z$ is dual to the discrete 'topological' time $n$, and $t_{p}(z, s, \beta)$ is the weight of the cycle $p$.
$A_{p}$ is the integrated observable $a(x)$ evaluated on a single traversal of cycle $p$

$$
A_{p}=\left\{\begin{array}{ll}
\int_{0}^{T_{p}} a\left(\boldsymbol{f}^{\tau}\left(\boldsymbol{x}_{0}\right)\right) \mathrm{d} \tau & \text { (flows) }  \tag{3}\\
\sum_{k=0}^{n_{p}-1} a\left(f^{k}\left(x_{0}\right)\right) & \text { (maps) }
\end{array} \quad x_{0} \in p\right.
$$

Here the functions $a(x)$ on $\mathcal{P}$ and $a(x)$ on $\mathcal{M}$ are two distinct functions, with $\boldsymbol{x}_{0}, x_{0}$ indicating that the observable is defined on the full flow, Poincaré surface of section, respectively. Later, we apply the same convention to the invariant densities $\rho$.

Classical averages over chaotic systems are given by cycle expansions constructed from derivatives of dynamical zeta functions. By expanding the product (2) a dynamical zeta function can be represented as a cycle expansion

$$
\begin{align*}
1 / \zeta & =1-\sum_{\pi}^{\prime} t_{\pi} \\
t_{\pi} & =t_{\pi}(z, s, \beta)=(-1)^{k_{\pi}+1} t_{p_{1}} t_{p_{2}} \ldots t_{p_{k}} \\
& =(-1)^{k_{\pi}+1} \frac{1}{\left|\Lambda_{\pi}\right|} \mathrm{e}^{\beta A_{\pi}-s T_{\pi}} z^{n_{\pi}} \tag{4}
\end{align*}
$$

where the prime on the sum indicates that the sum is restricted to pseudocycles, built from all distinct products of non-repeating prime cycle weights. The pseudocycle topological length, period, integrated observable, and stability are

$$
\begin{array}{ll}
n_{\pi}=n_{p_{1}}+\cdots+n_{p_{k}} & T_{\pi}=T_{p_{1}}+\cdots+T_{p_{k}}  \tag{5}\\
A_{\pi}=A_{p_{1}}+\ldots+A_{p_{k}} & \Lambda_{\pi}=\Lambda_{p_{1}} \Lambda_{p_{2}} \cdots \Lambda_{p_{k}}
\end{array}
$$

For economy of notation we usually omit the explicit dependence of $1 / \zeta$ and $t_{p}$ on $(z, s, \beta)$ whenever the dependence is clear from the context.

Truncation of the dynamical zeta function with respect to the topological length $n_{\pi} \leqslant N$ is indicated by a subscript:

$$
\begin{equation*}
1 / \zeta_{(N)}(z, s, \beta)=1-\sum_{n_{\pi} \leqslant N}^{\prime} t_{\pi} . \tag{6}
\end{equation*}
$$

If the system is bounded (such that no trajectories escape), the dynamical zeta function (2) has a leading zero at $1 / \zeta(1,0,0)=0$. Expressing this condition in terms of the cycle expansion (4) we find that any bound system satisfies the flow conservation sum rule [3]:

$$
\begin{equation*}
1 / \zeta(1,0,0)=1-\sum_{\pi}^{\prime}(-1)^{k_{\pi}+1} \frac{1}{\left|\Lambda_{\pi}\right|}=0 \tag{7}
\end{equation*}
$$

If the dynamics is ergodic, and the observable is sufficiently regular, the cycle expansions for the phase space average of observable $a(x)$ are given by either the integral over the natural measure, or by the cycle expansions

$$
\begin{align*}
& \text { flows: }\langle a\rangle_{\text {flow }}=\int_{\mathcal{M}} a(\boldsymbol{x}) \rho(\boldsymbol{x}) \mathrm{d} \boldsymbol{x}=\frac{\langle A\rangle_{\zeta}}{\langle T\rangle_{\zeta}}  \tag{8}\\
& \text { maps: }\langle a\rangle_{\text {map }}=\int_{\mathcal{P}} a(x) \rho(x) \mathrm{d} x=\frac{\langle A\rangle_{\zeta}}{\langle n\rangle_{\zeta}} \tag{9}
\end{align*}
$$

Here $\rho(\boldsymbol{x})$ and $\rho(x)$ denote the the natural invariant densities on $\mathcal{M}$ and $\mathcal{P}$, respectively. As we show in (17), the averages computed from the two representations of dynamics are related by the mean free-flight time.

The cycle expansions required for the evaluation of periodic orbit averages (8) and (9) are given by derivatives of the dynamical zeta function with respect to $\beta, s$ and $z$ :

$$
\begin{align*}
& \langle A\rangle_{\zeta}=\frac{\partial}{\partial \beta} 1 / \zeta(1,0,0)=\sum_{\pi}^{\prime}(-1)^{k_{\pi}+1} A_{\pi} /\left|\Lambda_{\pi}\right|  \tag{10}\\
& \langle T\rangle_{\zeta}=-\frac{\partial}{\partial s} 1 / \zeta(1,0,0)=\sum_{\pi}^{\prime}(-1)^{k_{\pi}+1} T_{\pi} /\left|\Lambda_{\pi}\right|  \tag{11}\\
& \langle n\rangle_{\zeta}=\frac{\partial}{\partial z} 1 / \zeta(1,0,0)=\sum_{\pi}^{\prime}(-1)^{k_{\pi}+1} n_{\pi} /\left|\Lambda_{\pi}\right| \tag{12}
\end{align*}
$$

Again, we use a subscript to indicate that the average is computed from a truncated zeta function, for instance

$$
\begin{equation*}
\langle A\rangle_{\zeta,(N)}=\frac{\partial}{\partial \beta} 1 / \zeta_{(N)}(1,0,0) \tag{13}
\end{equation*}
$$

## 3. Periodic orbit sum rules for billiards

We start by reviewing the mean free-flight time sum rule for billiards discussed by Chernov in [14].

In a $d$-dimensional billiard, a point particle moves freely inside a domain $Q$, scattering elastically off its boundary $\partial Q$. The billiard flow $f^{t}$ on $\mathcal{M}=Q \times S^{d-1}$ (where $S$ is the unit sphere of velocity vectors) has a natural Poincaré surface of section associated with the boundary

$$
\begin{equation*}
\mathcal{P}=\partial \mathcal{M}=\{(q, v) \in \mathcal{M}: q \in \partial Q \text { and } v \cdot n(q) \geqslant 0\} \tag{14}
\end{equation*}
$$

where $n(q)$ is the inward normal vector to the boundary at $q$, defined everywhere except at the singular set $\partial \mathcal{M}^{*}$ of non-differentiable points of the boundary (such as corners and cusps). In what follows we restrict the discussion to two-dimensional billiards.

Assume that the particle has unit mass and moves with unit velocity, $p_{1}^{2}+p_{2}^{2}=1$. The Cartesian coordinates and their conjugate momenta for the full phase space, $\mathcal{M}$, of the billiard are

$$
\boldsymbol{x}=\left(q_{1}, q_{2}, p_{1}, p_{2}\right)=\left(q_{1}, q_{2}, \sin \phi, \cos \phi\right)
$$

Let the Poincaré map be the boundary-boundary map $f: \partial \mathcal{M} \rightarrow \partial \mathcal{M}$, and parametrize the boundary $\partial \mathcal{M}$ by the Birkhoff (area preserving) coordinates

$$
x=\left(s, p_{s}\right) \quad p_{s}=\sin \theta
$$

where $s$ is the arc length measured along the boundary, $\theta$ is the scattering angle measured from the outgoing normal, and $p_{s}$ is the component of the momentum parallel to the boundary. Both the area of the billiard $A=|Q|$ and its perimeter length $L=|\partial Q|$ are assumed finite. Let $\tau(x)$ be the time of flight until the next collision. The continuous trajectory is parametrized by the Birkhoff coordinates together with a time coordinate $0<t<\tau(x)$ measured along the ray emanating from the boundary point $x=\left(s, p_{s}\right)$.

The period of a cycle $p$ is the sum of the finite free-flight segments

$$
T_{p}=\sum_{k=0}^{n_{p}-1} \tau\left(f^{k}\left(x_{0}\right)\right)
$$

where $x_{0}=\left(s_{0}, p_{s_{0}}\right)$ is any of the collision points in cycle $p$. The mean free-flight time is the average time of flight between successive bounces off the billiard boundary. It can be expressed either as a time average

$$
\bar{\tau}\left(x_{0}\right)=\lim _{n \rightarrow \infty} \frac{1}{n} \sum_{i=0}^{n-1} \tau\left(f^{i}\left(x_{0}\right)\right)
$$

or, as a phase space average

$$
\begin{equation*}
\langle\tau\rangle=\int_{\mathcal{P}} \tau(x) \mathrm{d} \mu(x) \tag{15}
\end{equation*}
$$

where $\mathrm{d} \mu(x)=\mathrm{d} s \mathrm{~d} p_{s} / \int_{\mathcal{P}} \mathrm{d} s \mathrm{~d} p_{s}$ is the natural measure. For Hamiltonian flows such as the billiard flow considered here this is simply the Lebesgue measure. If the billiard is ergodic, the time average is defined and independent of $x_{0}$ for almost all $x_{0}$. In order to find an exact expression for the phase space average $\langle\tau\rangle$, compute the integral over the entire phase space of the billiard,

$$
\int_{\mathcal{M}} \delta\left(1-p_{1}^{2}-p_{2}^{2}\right) \mathrm{d} q_{1} \mathrm{~d} q_{2} \mathrm{~d} p_{1} \mathrm{~d} p_{2}=2 \pi A
$$

and recompute the same thing in the Birkhoff coordinates,

$$
\begin{align*}
\int \delta\left(1-p_{1}^{2}-p_{2}^{2}\right) \mathrm{d} q_{1} \mathrm{~d} q_{2} \mathrm{~d} p_{1} \mathrm{~d} p_{2} & =\int_{\mathcal{P}} \mathrm{d} s \mathrm{~d} p_{s} \int_{t=0}^{\tau(x)} \mathrm{d} t=\int_{\mathcal{P}} \tau(x) \mathrm{d} s \mathrm{~d} p_{s} \\
& =\langle\tau\rangle \int_{\mathcal{P}} \mathrm{d} s \mathrm{~d} p_{s}=2 L\langle\tau\rangle \tag{16}
\end{align*}
$$

where $L$ is the circumference of the billiard. Hence the mean free-flight time is a purely geometric property of the billiard,

$$
\begin{equation*}
\langle\tau\rangle=\pi \frac{A}{L} \tag{17}
\end{equation*}
$$

the ratio of its perimeter to its area. The relation is a consequence of the explicit form of the invariant measure and applies to any billiard regardless of whether its phase space is mixed or not. For ergodic systems the periodic orbit theory gives a cycle expansion formula (9) for the mean free-flight time

$$
\begin{equation*}
\langle\tau\rangle=\frac{\langle T\rangle_{\zeta}}{\langle n\rangle_{\zeta}} \tag{18}
\end{equation*}
$$

If we know $\langle\tau\rangle$ this formula enables us to relate any discrete time average (9) computed from the map to the continuous time averages (8) computed on the flow. They are linked by the mean free-flight time formula

$$
\begin{equation*}
\langle a\rangle_{m a p}=\langle a\rangle_{\text {flow }}\langle\tau\rangle . \tag{19}
\end{equation*}
$$

As the next example of a periodic orbit sum rule, consider the case of the observable being the transverse momentum change at collision, $a=2 \cos \theta$. The corresponding sum rule is called the pressure sum rule because it is related to the pressure exerted by the particle on the billiard boundary.

The average pressure is given by the relation $P=F /|\partial Q|$, where $F$ is the time average of momentum change, that is the force the particle exerts against the boundary. The momentum change per bounce equals twice the transverse momentum at the collision, so the average force per bounce is
$\langle F\rangle_{\text {map }}=\int_{\mathcal{P}} 2 p_{\perp}(x) \mathrm{d} \mu(x)=\frac{1}{\int_{\mathcal{P}} \mathrm{d} s \mathrm{~d} p_{s}} \int_{\partial Q} \int_{-\pi / 2}^{\pi / 2} 2 \cos \theta \cos \theta \mathrm{~d} \theta \mathrm{~d} s=\frac{\pi}{2}$.
Hence the pressure for a flow becomes $\dagger$

$$
\begin{equation*}
\langle P\rangle_{\text {flow }}=\frac{\langle P\rangle_{\text {map }}}{\langle\tau\rangle}=\frac{\langle F\rangle_{\text {map }}}{L\langle\tau\rangle}=\frac{1}{2 A} . \tag{21}
\end{equation*}
$$

The exact averages (17), (21) apply to billiards of any shape, ergodic or not. As for ergodic billiards, both the mean free path and pressure can be calculated by means of cycle expansions, these relations lead to exact billiard sum rules for ergodic systems.

In what follows we restrict our attention to map averages, and omit the subscript $\langle\ldots\rangle_{\text {map }} \rightarrow\langle\ldots\rangle$.

Any average of an observable $a(x)$, defined in terms of $\partial \mathcal{M}$ coordinates $x=\left(s, p_{s}\right)$ can be expressed in terms of a simple integral since the invariant measure $\mathrm{d} \mu(x)=\rho(x) \mathrm{d} x=$ $\mathrm{d} s \mathrm{~d} p_{s} / \int_{\mathcal{P}} \mathrm{d} s \mathrm{~d} p_{s}$ is known explicitly. For each such observable we obtain an exact periodic orbit sum rule

$$
\begin{equation*}
\langle a\rangle=\int_{\mathcal{P}} a(x) \mathrm{d} \mu(x)=\frac{\langle A\rangle_{\zeta}}{\langle n\rangle_{\zeta}} \tag{22}
\end{equation*}
$$

$\dagger$ We are grateful to C P Dettmann and G P Morriss for the derivation of this formula (unpublished).

The integral defines the average $\langle a\rangle$, with no assumptions about ergodicity. If, in addition, the system is ergodic, by (8) we obtain an exact cycle expansion sum rule.

This uncountable infinity of sum rules seems not to have been noted in the literature.
The formula (22) does not allow for analytical computation of every average we want to compute in a billiard. Consider the simplest nontrivial average worthy of study in billiards, the diffusion coefficient

$$
\begin{equation*}
\langle D\rangle=\left.\frac{1}{2 d} \frac{1}{\langle n\rangle_{\zeta}} \frac{\partial^{2}}{\partial \beta^{2}}(1 / \zeta)\right|_{\beta=0} . \tag{23}
\end{equation*}
$$

As this formula requires evaluation of a second derivative of the relevant dynamical zeta function, the two-point correlations of the observable along cycles will enter the averaging formulae, and the average cannot be computed from one iterate of the map.

Another quantity of interest is the Lyapunov exponent. Let $\Lambda\left(x_{0}, n\right)$ be the largest eigenvalue of the Jacobian of the $n$th iterate of the map. The (largest) Lyapunov exponent is defined as

$$
\lambda=\lim _{n \rightarrow \infty} \frac{1}{n} \log \left|\Lambda\left(x_{0}, n\right)\right| .
$$

The cycle expansion formulae in section 2 compute

$$
\begin{equation*}
\lambda=\lim _{n \rightarrow \infty} \frac{1}{n} \int_{\mathcal{P}} \log |\Lambda(x, n)| \rho(x) \mathrm{d} x \tag{24}
\end{equation*}
$$

that is, a combination of time and phase space averages. Note that if $\Lambda\left(x_{0}, n\right)$ is multiplicative, $\Lambda\left(x_{0}, n\right)=\prod_{k=0}^{n-1} \Lambda\left(f^{k}\left(x_{0}\right)\right)$, as is the case for one-dimensional maps, then the integral in (24) is independent of $n$. In particular, we can set $n=1$ and reduce the problem to one iterate of the map. However, in most cases the invariant measure $\rho(x)$ is not known a priori, and there is no simple exact formula for the average.

For billiards the problem is the reverse: the invariant density is known but the expanding stability eigenvalue $\Lambda\left(x_{0}, n\right)$ is not multiplicative along an arbitrary trajectory, and the integral in (24) is dependent on $n$. It is possible to derive a multiplicative evolution operator for this purpose [15]. However, for the purpose at hand the naive cycle expansion formulae still apply, because $\Lambda\left(x_{0}, n\right)$ is multiplicative for repeats of periodic orbits. By defining the cycle weight

$$
\mathrm{e}^{\beta A_{p}}=\left|\Lambda_{p}\right|^{\beta}
$$

the cycle expansion for the Lyapunov exponent is given by

$$
\begin{equation*}
\langle\lambda\rangle=\frac{\langle\ln | \Lambda| \rangle_{\zeta}}{\langle n\rangle_{\zeta}} . \tag{25}
\end{equation*}
$$

So even though Lyapunov only requires computation of two first-order derivatives of the dynamical zeta function, it requires $n$-point correlations to all orders and cannot be computed by a sum rule.

In the case when (19) relates the Lyapunov exponent of the flow to the Lyapunov exponent of the corresponding Poincaré return map, the relation first proven by Abramov [16].

## 4. The overlapping three-disc billiard

We will test the above sum rules on cycle expansions for a concrete system, the overlapping three-disc billiard. This billiard consists of three discs of radius $a$ centred on the corners of an equilateral triangle with sides $R$. There is a finite enclosure (see figure 1) between the discs if $\sqrt{3}<R / a \leqslant 2$. This enclosure defines the billiard domain $Q$. One of the limits $R / a \rightarrow \sqrt{3}$ corresponds to the integrable equilateral triangle billiard. The other limiting case $R / a=2$


Figure 1. The overlapping three-disc billiard. A point-like particle moves inside the billiard bouncing specularly off the boundary. Shown is a cycle of topological length 4.

Table 1. The mean free-flight time $\langle\tau\rangle$, the average pressure $P$, and the best estimate of the Lyapunov exponent $\lambda$ computed by cycle expansion as function of the three-disc centre-to-centre separation $R$ used in our numerical tests, with disc radius fixed to $a=1$. For $R=1.9$ a numerically computed Lyapunov reference value obtained by direct simulation using $10^{10}$ bounces is displayed. The total numbers of the fundamental domain prime cycles used in the cycle expansion computations are also indicated.

| $R$ | $\langle\tau\rangle$ | $\langle\lambda\rangle$ | $\lambda_{\text {num }}$ | $L\langle P\rangle$ | \# cycles |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 1.85 | 0.102 | 0.523 |  | 1.570 | 342 |
| 1.90 | 0.1401 | 0.6036 | 0.60363 | 1.570 | 525 |

exhibits intermittency with infinite sequences of periodic orbits whose periods $T_{p}$ accumulate to finite limits, and where stabilities fall off as some power $n_{p}^{\alpha}$, where $n_{p}$ is the topological length.

The $\mathrm{C}_{3 v}$ symmetry of the billiard enables us to work in the fundamental domain [17]. The fundamental domain is a one-sixth slice of the billiard domain, fenced in by the symmetry lines of the billiard. In what follows we are only interested in the lowest eigenvalue and therefore we restrict our computations to the fully symmetric $A_{1}$ subspace. The fundamental domain symbolic dynamics is binary, but is not of the finite subshift type; its full specification would require an infinity of pruning rules of arbitrary length.

The mean free-flight time (17) for the overlapping three-disc billiard can be found by geometric considerations:

$$
\begin{equation*}
\langle\tau\rangle=\pi \frac{R^{2} / 4 \sqrt{3}-a^{2} \theta-R r / 2}{2 a \theta} \tag{26}
\end{equation*}
$$

where $r=\sqrt{a^{2}-(R / 2)^{2}}$ and $\theta=\pi / 6-\arcsin (r / a)$. We shall set $a=1$ throughout this paper, and parametrize the billiard by the centre-to-centre distance $R$. All our numerical tests are done for $R=1.9$. Results for this parameter value, as well as for $R=1.85$ are shown in table 1.

Figure 2 illustrates the convergence of finite topological length cycle expansions for the flow conservation sum rule (7) and for the mean free-flight time sum rule (18). As the exact result is known, we plot the logarithm of the error as a function of the truncation length $N$.

The overall exponential convergence indicates the existence of gaps in the zeta function. This would mean that $1 / \zeta(z, 0,0)$ is analytic and free of zeros in a disc extending beyond $z=1$ and $1 / \zeta(1, s, 0)$ is analytic and free of zeros in a halfplane extending beyond $s=0$. In the following we work under the milder assumption of only analyticity in the disc and the halfplane respectively: we will refer to this as assumption A. The 'irregular' oscillations in figure 2 are typical for systems with complicated symbolic dynamics and may reflect the


Figure 2. Convergence of cycle expansions: deviation of cycle expansions truncated to the topological length $N$ from exact sum rules for (o) flow conservation (7) and ( $\square$ ) mean free-flight time (18).
existence of a natural boundary [18].
For systems with finite-subshift symbolic dynamics the oscillations cease when the cutoff, $N$, exceeds the longest forbidden substring and, if the full spectral determinant is used, superexponential convergence $[5,13]$ sets in.

One should note the coincidence of the peaks and dips of the two curves. This type of correlation between coefficients of different power series will be important in the following.

## 5. Utilizing exact sum rules

We illustrate the utility of exact sum rules in accelerating the convergence of cycle expansions by applying the flow conservation sum rule to the problem of computing the mean free-flight time (17). As we already have the exact formula for this average, we are able to compute the exact error in the various estimates and compare them. We then apply the same technique to evaluate the Lyapunov exponent, for which no exact formula exists.

The traditional estimate uses truncated zeta functions in a straightforward fashion:

$$
\begin{equation*}
\langle\tau\rangle_{(N)}=\frac{\langle T\rangle_{\zeta,(N)}}{\langle n\rangle_{\zeta,(N)}} . \tag{27}
\end{equation*}
$$

The idea is to use the flow conservation sum rule (7) to improve the numerator and denominator of (27) separately. We begin with the denominator $\langle n\rangle_{\zeta}=\frac{\partial}{\partial z} 1 / \zeta(z=1,0,0)$.

The general problem is to find a good estimate of the derivative $F^{\prime}\left(z_{0}\right)$ at the first root $F\left(z_{0}\right)=0$ of a function given by a power series $F(z)=\sum_{k=0}^{\infty} b_{k} z^{k}$, given only a truncated version of the function:

$$
\begin{equation*}
F_{(N)}(z)=\sum_{k=0}^{N} b_{k} z^{k} \tag{28}
\end{equation*}
$$

In the appendix we show that, under assumption $A$ in section 4, such an estimate is given by

$$
\begin{equation*}
F^{\prime}\left(z_{0}\right) \approx F_{(N)}^{\prime}\left(z_{0}\right)-\frac{N+1}{z_{0}} F_{(N)}\left(z_{0}\right) \tag{29}
\end{equation*}
$$

In the appendix we argue that the error in the above estimate is suppressed compared with the error of the estimate $F_{(N)}^{\prime}\left(z_{0}\right)$ by a factor $q$ whose asymptotic behaviour is

$$
\begin{equation*}
q \sim 1 / N . \tag{30}
\end{equation*}
$$

The traditional estimate of $\langle n\rangle_{\zeta}$ was $\langle n\rangle_{\zeta,(N)}=F_{(N)}^{\prime}\left(z_{0}=1\right)$ and an improved estimate is now given by

$$
\begin{equation*}
\langle n\rangle_{\zeta, a c c} \equiv\langle n\rangle_{\zeta,(N)}-(N+1) \zeta_{(N)}^{-1}(1,0,0) . \tag{31}
\end{equation*}
$$

We proceed with the denominator $\langle T\rangle_{\zeta}=-\frac{\partial}{\partial s} 1 / \zeta(1, s=0,0)$. The problem is similar to the previous one but $1 / \zeta(1, s, 0)=0$ is now a Dirichlet series in $s$. According to assumption A, the zeta function $1 / \zeta(1, s, 0)$ is analytic in the halfplane $\operatorname{Re}(s)>0$ and the basic idea is to make a Taylor expansion of the zeta function

$$
\begin{equation*}
1 / \zeta(1, s, 0)=1-\sum_{\pi}^{\prime} \frac{(-1)^{k_{\pi}-1}}{\left|\Lambda_{\pi}\right|} \mathrm{e}^{-s T_{\pi}} \tag{32}
\end{equation*}
$$

around some point $s_{0}\left(\operatorname{Re}\left(s_{0}\right)>0\right)$
$1 / \zeta(1, s, 0)=1-\sum^{\prime} \frac{(-1)^{k_{\pi}-1}}{\left|\Lambda_{\pi}\right|} \mathrm{e}^{-s_{0} T_{\pi}} \sum_{k=0}^{\infty} \frac{\left(s_{0}-s\right)^{k} T_{\pi}^{k}}{k!} \equiv \sum_{k=0}^{\infty} c_{k}\left(s_{0}-s\right)^{k}$
where

$$
\begin{equation*}
c_{k}=\delta_{0, k}-\sum_{\pi}^{\prime}(-1)^{k_{\pi}-1} \frac{\mathrm{e}^{-s_{0} T_{\pi}} T_{\pi}^{k}}{\left|\Lambda_{\pi}\right| k!} . \tag{34}
\end{equation*}
$$

Convergence of the sum (32) implies convergence of the coefficients (34) (as can be realized from elementary properties of Dirichlet series) and can be identified with the coefficients of the desired Taylor series.

All pseudo-orbit sums are still truncated with respect to topological length. The coefficients $c_{k,(N)}$ can only be expected to approximate $c_{k}$ well for small enough $k$. Therefore, we truncate the series:

$$
\begin{align*}
1 / \zeta_{(N, M)}(1, s, 0) & =1-\sum_{\pi: n_{\pi} \leqslant N}^{\prime} \frac{(-1)^{k_{\pi}-1}}{\left|\Lambda_{\pi}\right|} \mathrm{e}^{-s_{0} T_{\pi}} \sum_{k=0}^{M} \frac{\left(s_{0}-s\right)^{k} T_{\pi}^{k}}{k!} \\
& \equiv \sum_{k=0}^{M} c_{k,(N)}\left(s_{0}-s\right)^{k} \tag{35}
\end{align*}
$$

where

$$
c_{k,(N)}=\delta_{0, k}-\sum_{\pi: n_{\pi} \leqslant N}^{\prime}(-1)^{k_{\pi}-1} \frac{\mathrm{e}^{-s_{0} T_{\pi}} T_{\pi}^{k}}{\left|\Lambda_{\pi}\right| k!} .
$$

This truncated Taylor series is the analogue of the truncated function $F_{(N)}(z)$ treated above; $s=0$ corresponds to $z=1$, and $s=s_{0}$ to $z=0$. From assumption A, we know that $1 / \zeta(1, s, 0)$ is analytic in a disc around $s=s_{0}$, extending beyond $s=0$, we can use (29) and derive the improved estimate:

$$
\begin{equation*}
\langle T\rangle_{\zeta, a c c} \equiv\langle T\rangle_{\zeta,(N)}+\frac{M+1}{s_{0}} \zeta_{(N, M)}^{-1}(1,0,0) . \tag{36}
\end{equation*}
$$

The choice of the maximal power $M$ depends on $s_{0}$, so how should $s_{0}$ and $M$ be chosen? Obviously, $s_{0}$ must lie somewhere in the range $1 / T_{\max }<s_{0}<1 / T_{\min }$ where the $T_{\min }$ and $T_{\max }$ are the smallest and largest period in the sample for a particular topological length cutoff $N$.

The next question is, for a given $s_{0}$, what is the number of reliable coefficients $c_{k,(N)}$ ? We see from (35) that pseudocycles are suppressed with their length according to the function $T^{k} \exp \left(-s_{0} T\right)$ having its maximum at $T=k / s_{0}$. So the coefficients with $k \ll s_{0} T_{\max }$ can


Figure 3. The error suppression factor for the improvement of the estimate of $\langle n\rangle_{\zeta}(\square)$ and for $\langle T\rangle_{\zeta}$ (o), applied to the three-disc system with $R=1.9$. Here we have used an extrapolated value from the cycle expansion as the best asymptotic estimate. Both error suppressions display the estimated error decrease and demonstrate that the sum rules do improve convergence.
be expected to be accurate. However, as the majority of cycles have periods close to $T_{\max }$ we want to make use of the information they carry. In our numerical work we have found it preferable to include a large number of fairly accurate coefficients rather than a small number of very accurate ones. So we choose the maximum power $M$ to be given by the average cycle length:

$$
M=\left.s_{0} \bar{T}_{p}\right|_{n_{p}=N-1}
$$

The error of the improved estimate is suppressed compared with the error of the traditional estimate by factor we call $q$, see the appendix. This $q$-factor is plotted in figure 3 . It decreases (apart from oscillations) as the estimated $N^{-1}$ error suppression derived for maps.

The calculation of the integrated observable amounts to evaluating the $\beta$ derivatives of the dynamical zeta functions. The role of $\beta$ is completely analogous to that of $s$. With $\beta$ viewed as a complex variable, the dynamical zeta function $1 / \zeta(1,0, \beta)$ is a Dirichlet series in $\beta$ and the above methods can be used to compute $\frac{\partial}{\partial \beta} 1 / \zeta(1,0,0)$. Here similar criteria apply to $\beta_{0}$ and $N$ as for (36): $\beta_{0}$ close to $1 / A_{\text {min }}$ and $M=\left.\beta_{0} \bar{A}_{p}\right|_{n_{p}=N-1}$.

### 5.1. Improvement on the averages

So far we have improved the numerator and the denominator of (18) and (9) separately. The errors of both are suppressed by a factor $q \approx \mathrm{O}(1 / N)$ compared with unaccelerated estimates. We have also seen (figure 3) that, both before and after resummation, their behaviour versus the cutoff $N$ are highly correlated. So it is not obvious how the resulting average should be improved, indeed it is not clear whether it is improved at all.

The accelerated cycle expansion for an observable $a(x)$ using our method is

$$
\begin{equation*}
\langle a\rangle_{a c c}=\frac{\langle A\rangle_{\zeta, a c c}}{\langle n\rangle_{\zeta, a c c}} \tag{37}
\end{equation*}
$$

and the error suppression, the $q$-factor for the observable $a(x)$ is

$$
\begin{equation*}
q_{a}=\frac{\langle a\rangle_{a c c}-\langle a\rangle_{e x a c t}}{\langle a\rangle_{(N)}-\langle a\rangle_{\text {exact }}} \tag{38}
\end{equation*}
$$

We compute this $q$ factor numerically for three different averages:
(i) The mean free-flight time $\langle\tau\rangle$ by (18).
(ii) The mean pressure $P$, cycle expansion of (21).
(iii) The Lyapunov exponent by (25). The reference value of the Lyapunov exponent is obtained by numerical simulation, see table 1 .


Figure 4. The error suppression factor (38) for: $(\diamond)$ the accelerated mean free-flight time sum rule, (o) pressure sum rule, $(\square)$ the Lyapunov exponent. At topological length 12, the accuracy of the accelerated Lyapunov exponent has reached the best estimate from direct numerical simulation (see table 1).

The results are summarized in figure 4. The accelerated cycle expansions are clearly better than the standard cycle expansions. The error suppression factors appear to decrease exponentially, and therefore the acceleration techniques has, for the three-disc system, increased the correlation between the expansions leading to a faster convergence for the averages.

## 6. Conclusion

In this paper we have achieved two objectives: (i) we have derived an infinite number of exact periodic orbit sum rules for billiards (22). Such sum rules enable us to make exact computations of some statistical averages for billiards, such as the mean free-flight time (17) and pressure (21). (ii) We have derived the improved estimate (36) which combines the flow conservation sum rule (7) with the cycle expansions. In order to measure the convergence acceleration, we have introduced the error suppression factor (38) that gauges the improvement of the accelerated cycle expansions relative to the unaccelerated ones. We thus demonstrate that exact sum rules can be used to accelerate convergence for observables for which no exact results exist, see figure 4.

A challenge for the future is to utilize such infinities of sum rules for billiards in the classical applications (other than the Lyapunov exponent studied here), as well as in the semi-classical applications of periodic orbit theory.

## Acknowledgments

This work was supported by the Swedish Natural Science Research Council (NFR) under contract nos F-AA/FU 06420-312 and F-AA/FU 06420-313. PD thanks NORDITA for partial support. SFN thanks PD and KTH for their hospitality.

## Appendix. Resummation of power series

Consider a function $F(z)=\sum_{k=0}^{\infty} b_{k} z^{k}$ given by a power series, where only a finite number of coefficients are known:

$$
\begin{equation*}
F_{(N)}(z)=\sum_{k=0}^{N} b_{k} z^{k} \tag{39}
\end{equation*}
$$

We assume that $F\left(z_{0}\right)=0$ for some $z_{0}$ and we wish to estimate the first derivative $F^{\prime}\left(z_{0}\right)$ (and possibly higher derivatives) as accurately as possible. A simple estimate is

$$
\begin{equation*}
F^{\prime}\left(z_{0}\right) \approx F_{(N)}^{\prime}\left(z_{0}\right) \tag{40}
\end{equation*}
$$

but this does not make use of the knowledge that $F\left(z_{0}\right)=0$.
In order to improve this estimate we follow the idea of [19]. We assume that the function is analytic in a disc enclosing $z_{0}$, cf assumption A . We can then consider a resummation of the Taylor series around $z=0$ into a Taylor series around $z=z_{0}$, from which we want to extract the desired coefficients. With only a finite number of coefficients at our disposal we make the ansatz

$$
\begin{equation*}
\sum_{k=0}^{N} b_{k} z^{k}=\sum_{i=1}^{N+1} a_{i}\left(z-z_{0}\right)^{i}+\mathrm{O}\left(z^{N+1}\right) \tag{41}
\end{equation*}
$$

where we have kept the number of known and unknown coefficients equal so that the system of equations is solvable. Note that the sum rule is built into this ansatz by setting $a_{0}=0$.

Later, we see that the estimate $F^{\prime}\left(z_{0}\right) \approx a_{1}$ is an improvement as compared with the simple estimate (40).

In order to compute $a_{1}$ we expand the right-hand side (41) in binomials

$$
\begin{equation*}
\sum_{i=0}^{N} b_{i} z^{i}=\sum_{i=1}^{N+1} a_{i} \sum_{j=0}^{i} z^{j}\left(-z_{0}\right)^{i-j}\binom{i}{j}+\mathrm{O}\left(z^{N+1}\right) \tag{42}
\end{equation*}
$$

we obtain the linear system of equations

$$
\begin{equation*}
b_{j}=\sum_{\max (j, 1)}^{N+1}\binom{i}{j}\left(-z_{0}\right)^{i-j} a_{i} \quad 0 \leqslant j \leqslant N \tag{43}
\end{equation*}
$$

To express this in a more convenient way we form $n$-dimensional vectors of $a_{i}$ and $b_{i}$

$$
\begin{equation*}
(\boldsymbol{b})_{i}=z_{0}^{i-1} b_{i-1} \quad(\boldsymbol{a})_{i}=z_{0}^{i} a_{i} \tag{44}
\end{equation*}
$$

where $n \equiv N+1$ and transfer to matrix notations and write (43) as

$$
\begin{equation*}
\boldsymbol{b}=\boldsymbol{M a} \quad(\boldsymbol{M})_{i j}=\binom{j}{i-1}(-1)^{j-i+1} \quad 1 \leqslant i, j \leqslant n . \tag{45}
\end{equation*}
$$

We use the convention that $\binom{n}{m}=0$ if $m$ is out of range. This system may readily be solved. Define the matrix $L$ by

$$
(\boldsymbol{L})_{i j}= \begin{cases}1 & i \geqslant j  \tag{46}\\ 0 & i<j .\end{cases}
$$

Then

$$
\begin{equation*}
(\boldsymbol{L} \boldsymbol{M})_{i j}=(-1)^{i+j+1}\binom{j-1}{i-1} \rightarrow(\boldsymbol{L} \boldsymbol{M})_{i j}^{-1}=-\binom{j-1}{i-1} \tag{47}
\end{equation*}
$$

and the explicit solution is

$$
\begin{equation*}
a=(L M)^{-1} L b \tag{48}
\end{equation*}
$$

In particular,

$$
\begin{equation*}
(\boldsymbol{a})_{1}=-n(\boldsymbol{b})_{1}-(n-1)(\boldsymbol{b})_{2}-\cdots-1(\boldsymbol{b})_{n}=-\sum_{k=1}^{n}(n-k+1)(\boldsymbol{b})_{k} \tag{49}
\end{equation*}
$$

which we rewrite in terms of our original notation:

$$
\begin{equation*}
z_{0} a_{1}=-\sum_{k=0}^{N}(N+1-k) z_{0}^{k} b_{k} \tag{50}
\end{equation*}
$$

The right-hand side is recognized as $z_{0} F_{(N)}^{\prime}\left(z_{0}\right)-(N+1) F_{(N)}\left(z_{0}\right)$, leading to our improved estimate

$$
\begin{equation*}
F^{\prime}\left(z_{0}\right) \approx a_{1}=F_{(N)}^{\prime}\left(z_{0}\right)-(N+1) z_{0}^{-1} F_{(N)}\left(z_{0}\right) \tag{51}
\end{equation*}
$$

The error is suppressed by a factor
$q=\frac{F_{N}^{\prime}\left(z_{0}\right)-(N+1) z_{0}^{-1} F_{N}\left(z_{0}\right)-F^{\prime}\left(z_{0}\right)}{F_{N}^{\prime}\left(z_{0}\right)-F^{\prime}\left(z_{0}\right)}=\frac{\sum_{k=N+1}^{\infty}(k-N-1) b_{k}}{\sum_{k=N+1}^{\infty} k b_{k}}$.
To express this in a more convenient form we use summation by parts, that is, we define

$$
\begin{equation*}
s_{k}=\sum_{j=k}^{\infty} b_{j} \tag{53}
\end{equation*}
$$

and write

$$
\begin{equation*}
1 / q=1+\frac{(N+1) s_{N}}{\sum_{k=N}^{\infty} s_{k}} \tag{54}
\end{equation*}
$$

If $F(z)$ is the spectral determinant for a $d$-dimensional Axiom A map, the coefficients of the power series expansion are super-exponentially bounded [5]:

$$
\begin{equation*}
C_{a} \Lambda_{a}^{-m^{1+1 / d}}<\left|b_{m}\right|<C_{b} \Lambda_{b}^{-m^{1+1 / d}} \tag{55}
\end{equation*}
$$

where $1<\Lambda_{b}<\Lambda_{a}$. Assuming, moreover, that the signs of the coefficients settle down to some periodic pattern, one can show that the error suppression factor has the following asymptotic behaviour:

$$
\begin{equation*}
q \sim N^{-(1+1 / d)} \tag{56}
\end{equation*}
$$

In this paper we focus on a hyperbolic systems whose symbolic dynamics cannot be finitely specified. However, if the zeta function is analytic in a disc enclosing $z_{0}$, cf assumption A, the bound on the coefficients is exponential

$$
\begin{equation*}
C_{a} \Lambda_{a}^{-m}<\left|b_{m}\right|<C_{b} \Lambda_{b}^{-m} \tag{57}
\end{equation*}
$$

and nothing can be said about the signs, as they can oscillate in a completely irregular fashion [18]. It seems difficult to obtain proper bounds on $q$ in a general setting. In the


Figure A1. Error suppression factor (38) ( $y$-axis) versus truncation in topological length $N$ ( $x$-axis) for the tent map (with a 'typical' slope value $\Lambda$ ).
case at hand we can only provide a qualified guess on the decrease of the error suppression factor:

$$
\begin{equation*}
q \sim N^{-1} \tag{58}
\end{equation*}
$$

Some evidence for this behaviour can be provided by the tent map

$$
F(x)= \begin{cases}\Lambda x & \text { for } \quad x<\frac{1}{2}  \tag{59}\\ -\Lambda(x-1) & \text { for } \quad x \geqslant \frac{1}{2}\end{cases}
$$

The expansion rate is uniform but complete symbolic dynamics is lacking in the generic case, that is for almost all $\Lambda$. It is then known [18] that the analyticity is limited by a natural boundary so this simple system can give us a hint of the behaviour of generic hyperbolic billiards. In figure A1 we plot the $q$-factor for the tent map for a randomly chosen parameter versus $N$. It conforms with the predicted $1 / N$ behaviour.

The ansatz (41) used here is the simplest conceivable and it led to very simple formulae. The only requirement is that the dynamical zeta function is analytic in a disc $z \leqslant R$, where $R>1$. This excludes strongly intermittent systems where a more refined ansatz is needed [19]. If one has some explicit knowledge of the nature of the leading singularity of the dynamical zeta function, one can tailor a more specific ansatz.

## References

[1] Gaspard P 1998 Chaos, Scattering and Statistical Mechanics (New York: Cambridge University Press)
[2] Gutzwiller M C 1990 Chaos in Classical and Quantum Mechanics (New York: Springer)
[3] Cvitanović P et al 1998 Classical and Quantum Chaos (Copenhagen: Niels Bohr Institute) webpage http://www.nbi.dk/ChaosBook
[4] Ruelle D 1976 Ergod. Theor. Dynam. Sys. 299
Baladi V 1995 Dynamical zeta functions Real and Complex Dynamical Systems ed B Branner and P Hjorth (Dordrecht: Kluwer)
[5] Rugh H H 1992 Nonlinearity 51237
[6] Cvitanović P, Rosenqvist P E, Rugh H H and Vattay G 1993 Chaos 3619
[7] Dettmann C P and Morriss G P 1997 Phys. Rev. Lett. 784201
[8] Dahlqvist P and Russberg G 1991 J. Phys. A: Math. Gen. 244763
[9] Hansen K T 1993 Symbolic dynamics in chaotic systems PhD Thesis University of Oslo webpage http://www.nbi.dk/CATS/papers/khansen
[10] Cvitanović P, Dettmann C T, Mainieri R and Vattay G 1998 J. Stat. Phys. 93981
[11] Cvitanović P, Rolf J, Hansen K and Vattay G 1998 Nonlinearity 111233
[12] Dahlqvist P 1995 Nonlinearity 811
[13] Artuso R, Aurell E and Cvitanović P 1990 Nonlinearity 3361
[14] Chernov N 1997 J. Stat. Phys. 881
[15] Cvitanović P and Vattay G 1993 Phys. Rev. Lett. 714138
[16] Abramov L M 1959 Dokl. Akad. Nauk 128647 Abramov L M 1959 Dokl. Akad. Nauk 128873
[17] Cvitanović P and Eckhardt B 1993 Nonlinearity 6277
[18] Dahlqvist P 1997 J. Math. Phys. 384273
[19] Dahlqvist P 1997 J. Phys. A: Math. Gen. 30 L351

